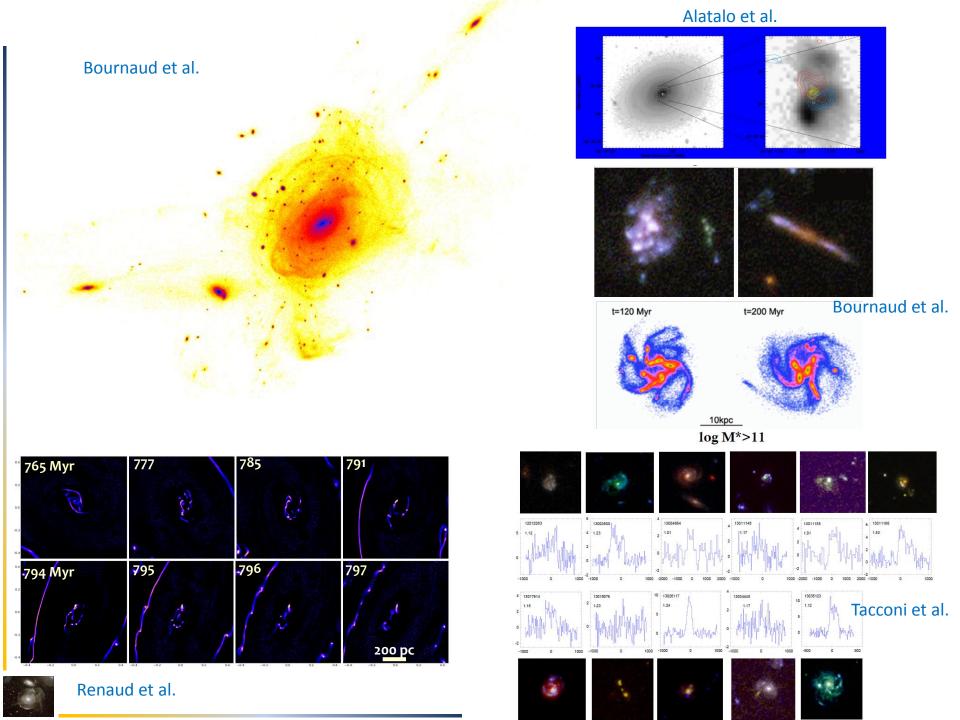
# **Extreme star forming environments** Near and far

Eric Emsellem



#### "Extreme": what do you mean?

- SFR, SFe ?
- High densities? Large mass (of H<sub>2</sub>?) in clumps?
- Large reservoirs? Depletion time-scale?
- Turbulent medium? Shocks? Pressure?
- Metallicity? Dust?
- So, are we talking about e.g. :
- Globular cluster formation
- Mergers, Strong Outflows
- Clumpy galaxies
- SM galaxies

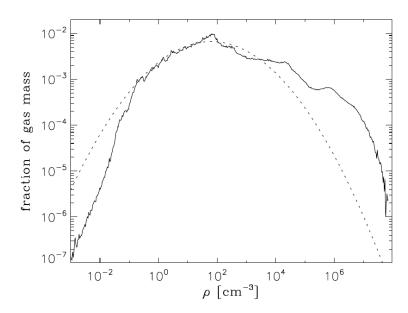


### "Extreme": what do you mean?

- SFR, SFe,
- High densities? Large mass (of H<sub>2</sub>?) in clumps?
- Large reservoirs? Depletion time-scale?
- Turbulent medium? Shocks? Pressure?
- Metallicity? Dust?

So, are we talking about e.g.:

- Globular cluster formation
- Clumpy galaxies
- Mergers, Strong Outflows
- SM galaxies



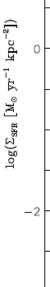
And why should these be special?

(is it just the PDF?)



Importance of the PDF?

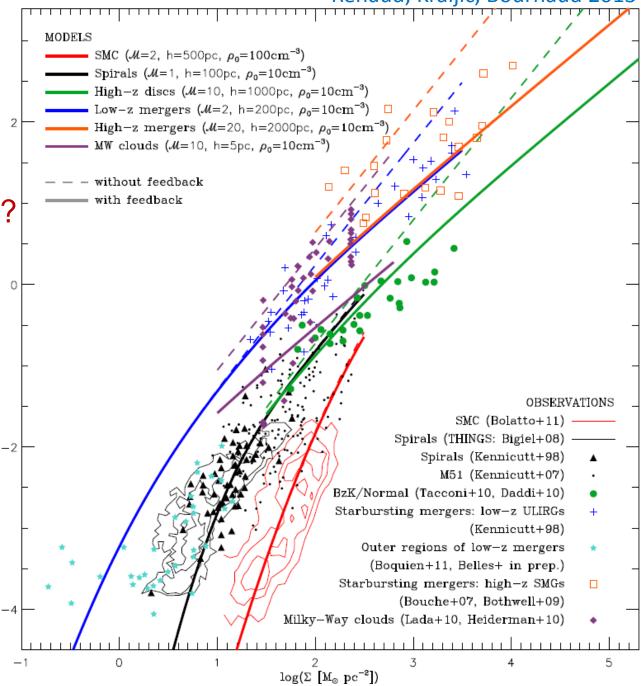
How do we get different PDFs?



Density

M

Self-gravity

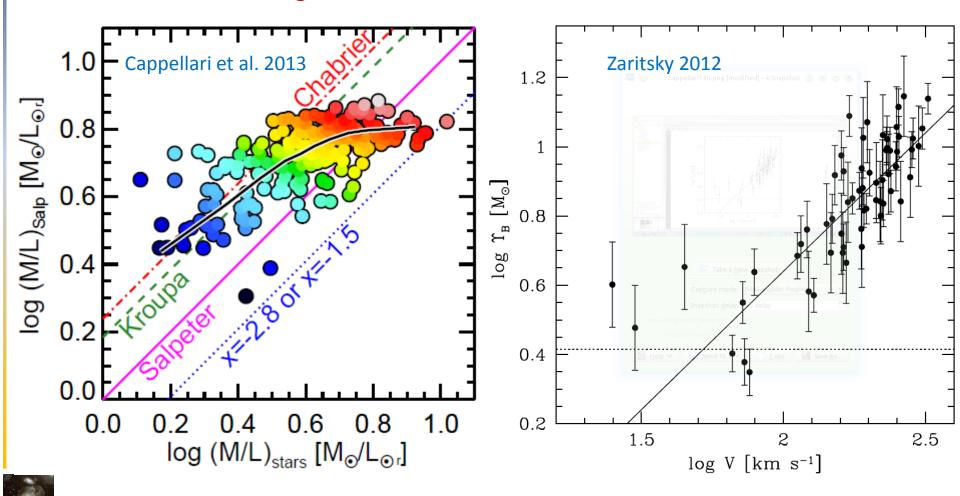


### Does feedback matter at all?

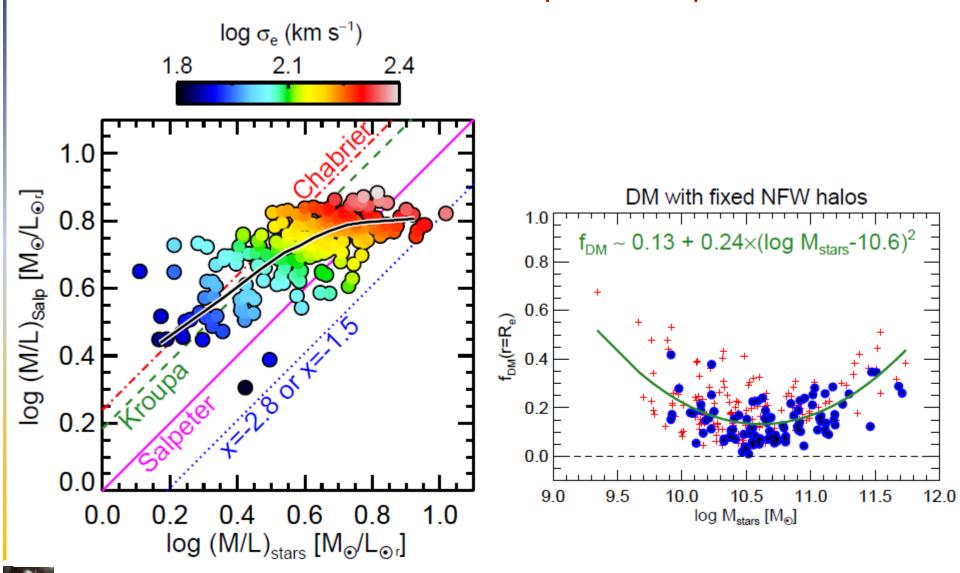
- Feeback HAPPENS but...
  - Cause or consequence ?
  - Self-)regulation ?

### IMF versus SF regime?

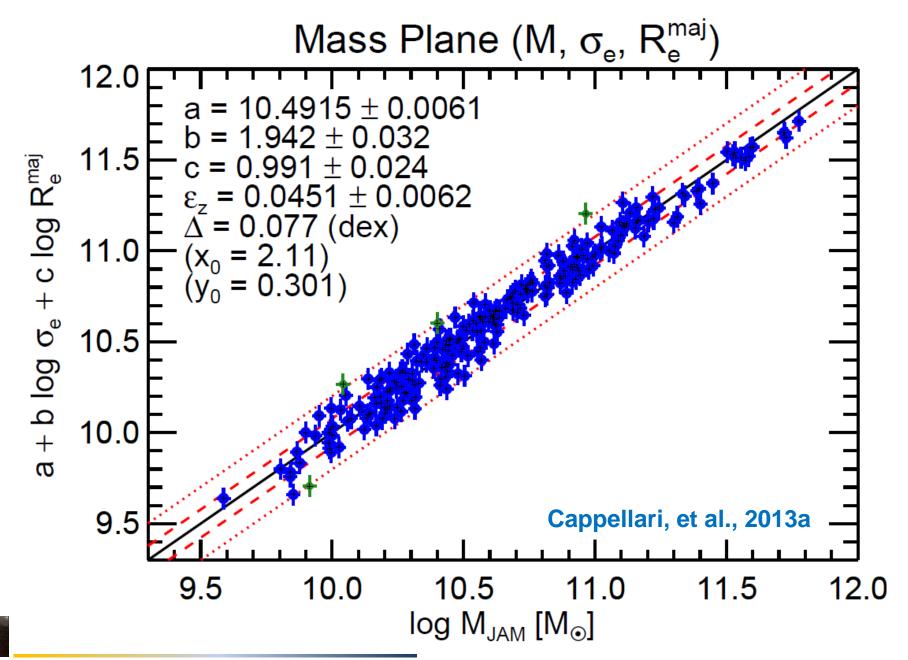
- Is there a link between SF regimes, environments and IMF?
- Is it the SF regime which determines the IMF or does IMF constrain SF regimes/environments?



- A link between these extreme environment and the IMF?
- DM fraction versus M similar to Specific frequencies of GCs?



## SF: what is your calibrator?



- Importance of physical processes
  - What are the important ingredients?
  - Is self-regulating or triggering important?
- Importance of environments
  - Is that defining the on-going physics?
  - Or is it just collateral damage?
- Link with Existing systems
  - Which tracers?
  - A How do we disentangle the ingredients?
- Perspectives?
  - IMF studies : how to decide if relevant
  - High redshift : how to connect with low z
  - Simulations: how to go beyond Zoology

