

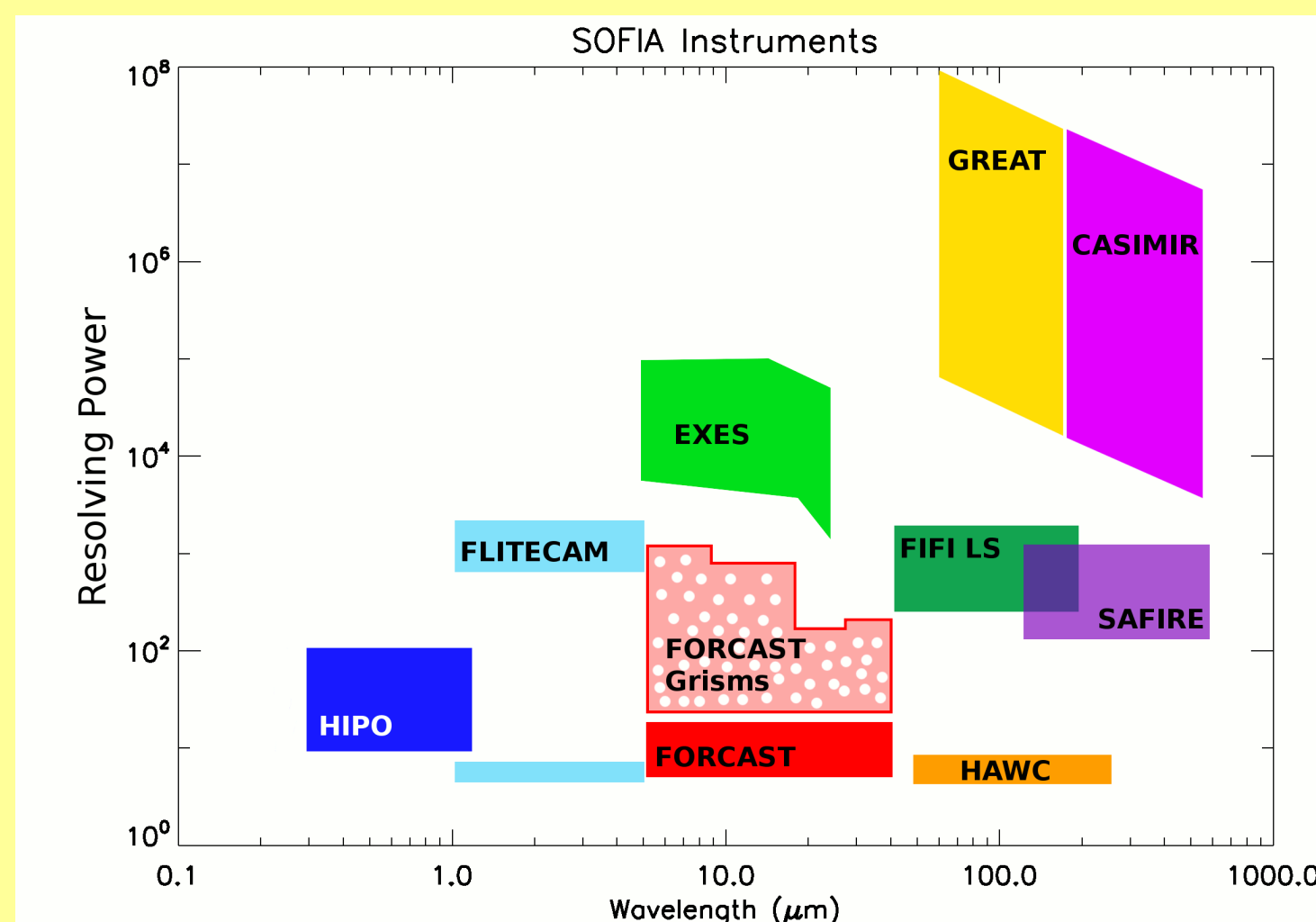
# A Silicon and KRS-5 Grism Suite for FORCAST on SOFIA

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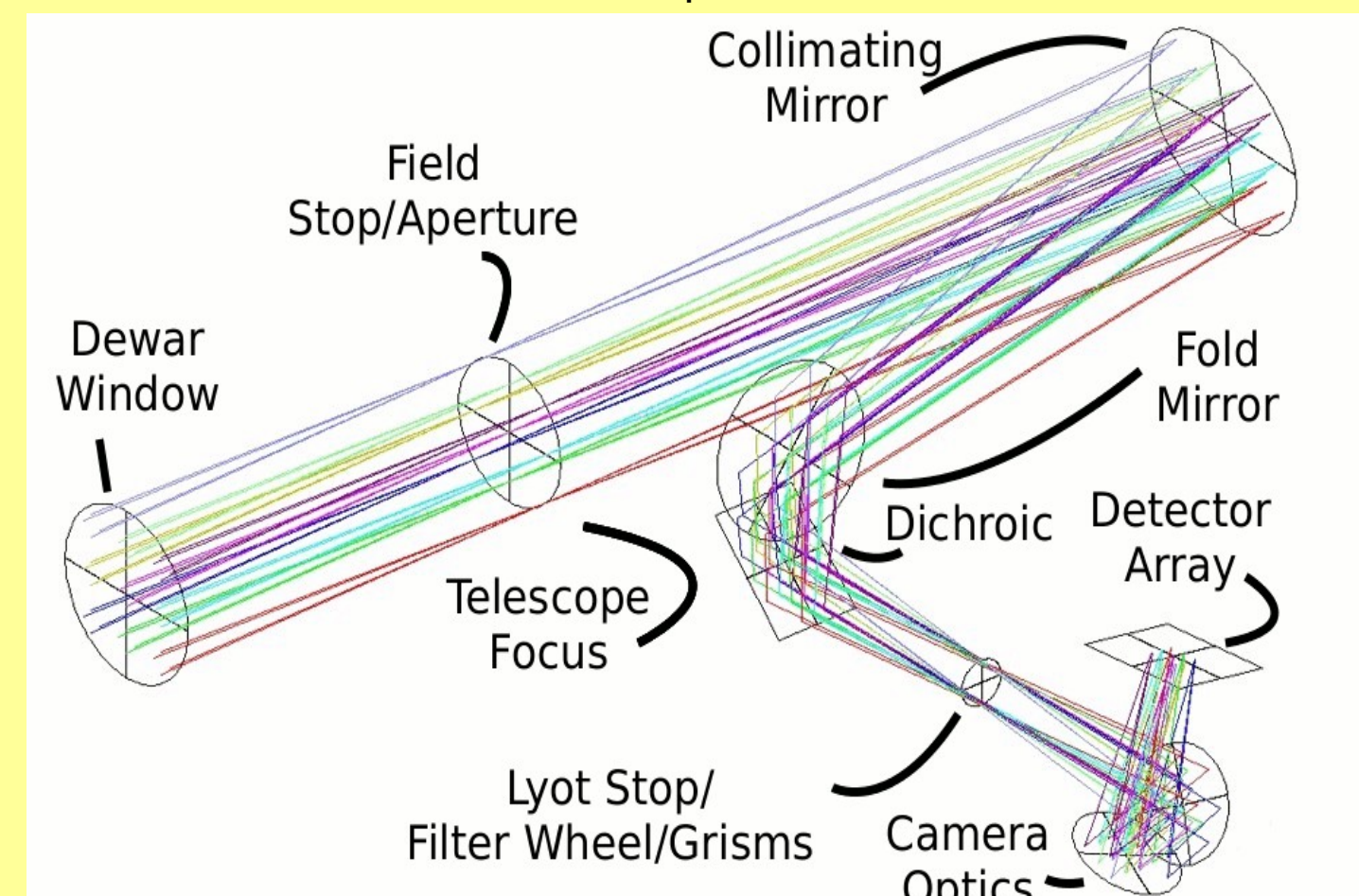
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## About FORCAST:

The Faint Object infraRed CAmera for the SOFIA Telescope (FORCAST) is a mid-infrared (5-40  $\mu\text{m}$ ) camera being built at Cornell University. While the sensitivity of Spitzer exceeds that of FORCAST, the angular resolution afforded by the 2.5 m telescope improves by a factor of  $\sim 3$  over that of Spitzer at  $\lambda > 15 \mu\text{m}$ . The camera provides a  $3.2' \times 3.2'$  field of view. Figure 2 shows the optical path layout of a single FORCAST channel.

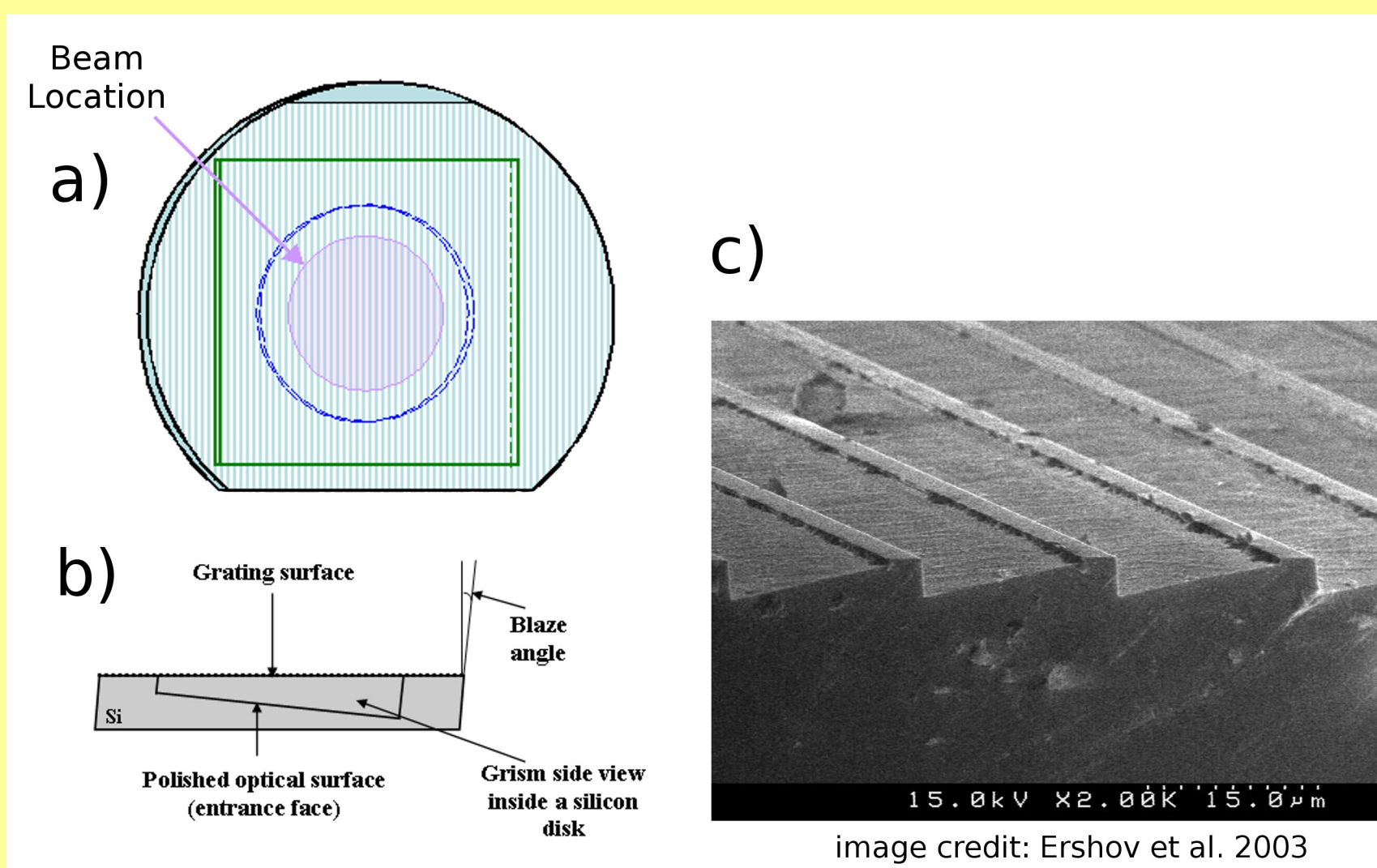


**Figure 1:** Wavelength vs. Resolving Power plot of the first generation of SOFIA instruments. One capability which the current SOFIA instrument complement lacks is mid-IR spectroscopy at moderate resolving power ( $R \sim 100-1000$ ). We report on the fabrication and testing of a suite of gratings for the FORCAST instrument which allows observers to use FORCAST as a low to medium resolution spectrometer.



**Figure 2:** Optical path of FORCAST short-wavelength channel. The gratings are mounted in filter wheels on either side of the Lyot stop.

## About Silicon Grisms:



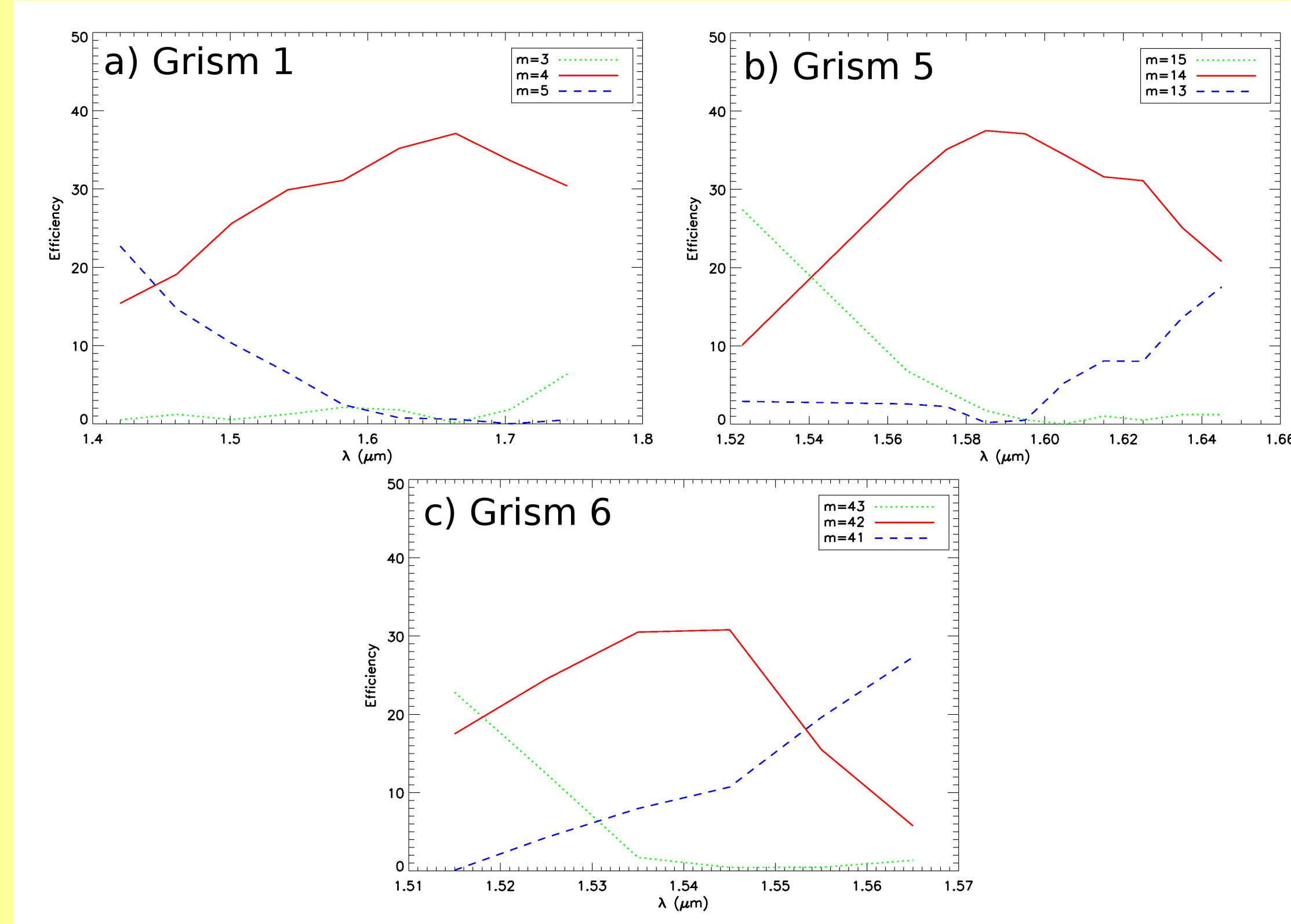
**Figure 3:** Grisms 1, 2, 5 and 6 were manufactured by chemically ruling grooves into silicon crystal. The details of this process can be found in Marsh et al. 2007 or Mar et al. (submitted to App Opt). In panel a, silicon disks are cut from the boule and polished. Thin lines are patterned onto the surface of the disk using lithographic techniques. During an anisotropic etch of silicon using a solution of KOH, grooves are "cut" along crystal planes, as shown in the SEM image of grooves in panel c. The grism is then cut from the silicon disk (panel b) and the entrance face is polished to optical flatness.

Grism	Material	Physical characteristics	Wavelength band	Grism size	Design R	Measured R	Measured efficiency at 1.4-1.8 $\mu\text{m}$
1	Si	$\sigma=25 \mu\text{m}$ , $\delta=6.16^\circ$ , $m=1$	5 - 8 $\mu\text{m}$	3'x2"	200	225	40 $\pm$ 3%
2	Si	$\sigma=87 \mu\text{m}$ , $\delta=32.6^\circ$ , $m=1$	5 - 8 $\mu\text{m}$	2"x15"	1200	1400	-
3	KRS-5	$\sigma=32 \mu\text{m}$ , $\delta=15.2^\circ$ , $m=1$	8.1 - 13.7 $\mu\text{m}$	3'x2"	300	-	-
4	KRS-5	$\sigma=130 \mu\text{m}$ , $\delta=36.8^\circ$ , $m=1$	8.1 - 13.7 $\mu\text{m}$	2"x15"	800	-	-
5	Si	$\sigma=87 \mu\text{m}$ , $\delta=6.16^\circ$ , $m=1$	17.1 - 28.1 $\mu\text{m}$	3'x2"	140	185	39 $\pm$ 3%
6	Si	$\sigma=142 \mu\text{m}$ , $\delta=11.07^\circ$ , $m=2$	28.7 - 40.0 $\mu\text{m}$	3'x3"	250	265	41 $\pm$ 3%

**Table 1.** FORCAST grism suite. The physical characteristics in column 3 are groove spacing ( $\sigma$ ), prism angle ( $\delta$ ), and order in which grism will operate ( $m$ ).

## Abstract

We have designed and fabricated a suite of gratings for use in FORCAST, a mid-infrared camera scheduled as a first light instrument on SOFIA. The grism suite gives SOFIA a new capability: low and moderate resolution spectroscopy from 5-37  $\mu\text{m}$ , without the addition of a new instrument. We fabricated four silicon ( $n = 3.44$ ) gratings using photolithographic techniques and purchased two additional mechanically ruled KRS-5 ( $n=2.3$ ) gratings. One pair of silicon gratings permits observations of the 5-8  $\mu\text{m}$  micron band with a long slit at resolving power ( $R$ ) of  $\sim 200$ , or in a cross-dispersed mode at  $R \sim 1200$ . In the 8 - 14  $\mu\text{m}$  region, where silicon absorbs heavily, the KRS-5 gratings will provide the spectroscopic capability at predicted resolving powers of 300 and 800 in long-slit and cross-dispersed mode, respectively. The remaining two silicon gratings cover 17 - 37  $\mu\text{m}$  at resolving powers of 140 and 250. We have thoroughly tested the silicon gratings in the laboratory, measuring efficiencies in transmission at 1.4 - 1.8  $\mu\text{m}$ . We report on these measurements as well as on cryogenic performance tests of the silicon and KRS-5 devices after installation in FORCAST.



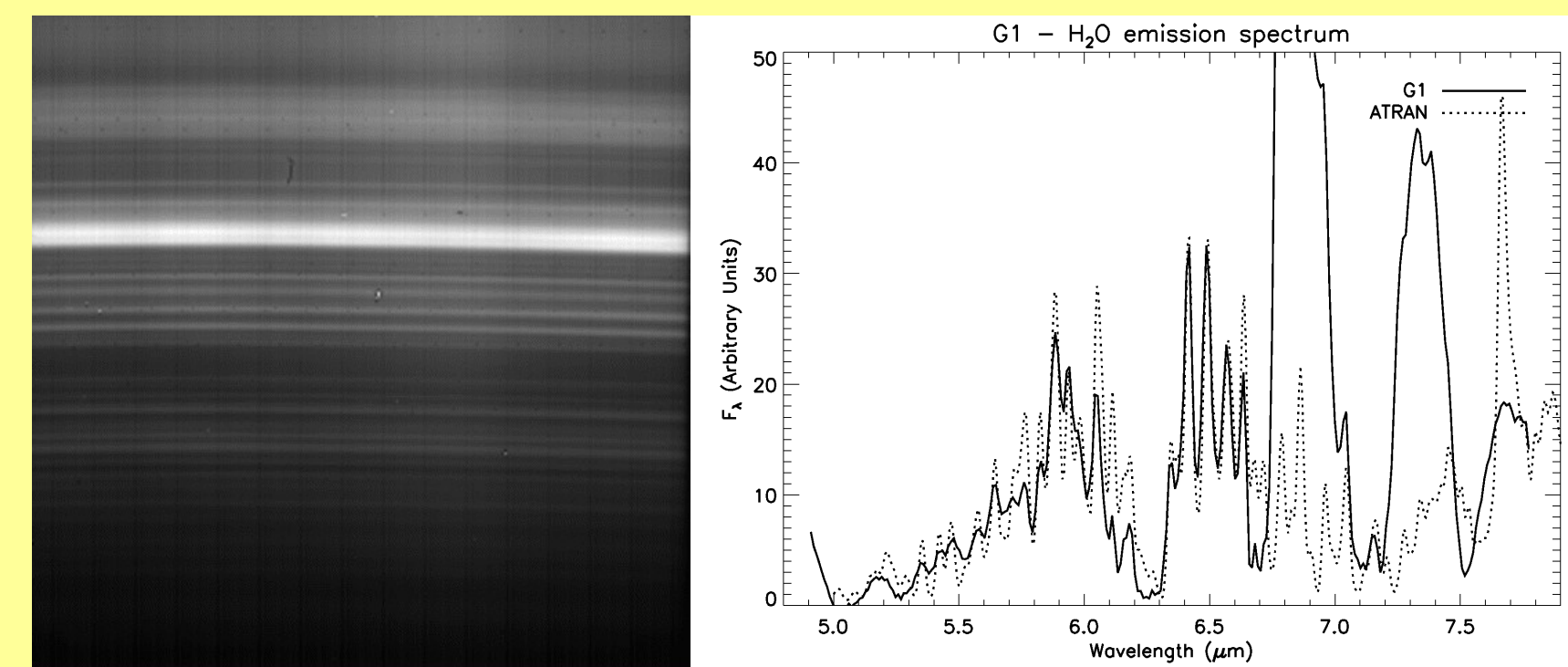
**Figure 4:** Measuring the blaze function of uncoated silicon gratings: Before installation in FORCAST, three out of four silicon gratings were tested using a scanning monochromator. We measured throughput efficiencies at several different wavelengths (1.4-1.8  $\mu\text{m}$ ) by collimating light from the output of a scanning monochromator into a 10mm diameter beam centered and normal to the grism entrance face and comparing it to the throughput of the system when no grism was present in the beam. The efficiency of the grism at a particular wavelength is then the intensity of the light integrated over several orders divided by the intensity measured with no grism in place. We did not measure the blaze function of grism 2 because the output of the monochromator was a significant fraction of the grism's free spectral range. Assuming 30% Fresnel losses at each surface and considering geometric losses appropriate to each grism, the measured efficiencies agree well with predicted efficiencies. The geometric losses are due to groove tops and groove shadowing, and are typically on the order of 10% for small wedge angles, but increase with the wedge angle.

In Figure 5, we show the results of FORCAST cooldown tests done in November 2006 and March 2008. Both silicon and KRS-5 gratings were fully integrated in FORCAST during these measurements. The two principal results are outlined below and in figures 5a-f on the right.

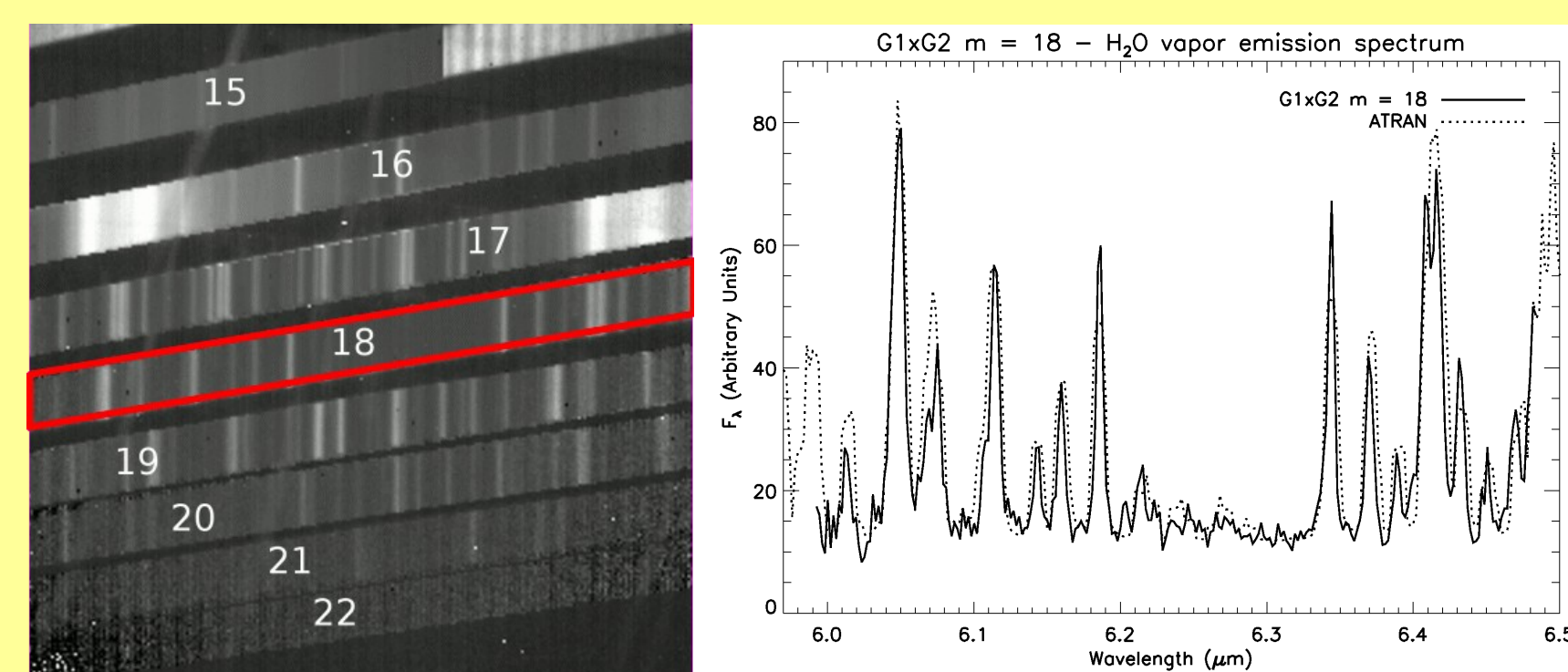
**Wavelength Calibration:** For the silicon gratings, one way to determine a wavelength calibration is to use water vapor emission lines. The wavelength calibration for G1, G1xG2, G5, and G6 was determined using the ATRAN model of telluric (mostly water vapor) lines smoothed to the specified resolution of each grism. A dewar of liquid nitrogen (LN2) was placed in the field of view of the camera. A path length of about 1 m of room-temperature water-vapor separated the camera window from the LN2 surface. The much warmer water vapor is seen in emission against the 77 K blackbody of the LN2. In the wavelength range of the KRS-5 gratings, (G3 and G3xG4, 8 - 13.7  $\mu\text{m}$ ) there are no significant water vapor lines to use as a wavelength calibration standard. Instead, we placed a hotplate at  $\sim 350\text{K}$  and placed a thin (1.5 mil) sheet of polystyrene across the camera window. The polystyrene has several broad absorption features in the spectral range of gratings 3 and 4 which served as wavelength calibration standards.

**Spectral Resolution:** Measuring the FWHM of unresolved water lines yielded resolutions of  $\sim 250$  and  $\sim 1400$  for gratings 1 and 2. Resolution measurements were not possible for gratings 3 and 4 since none of the polystyrene absorption lines are narrow enough to be unresolved. Gratings 5 and 6 show resolving powers of 185 and 265, respectively.

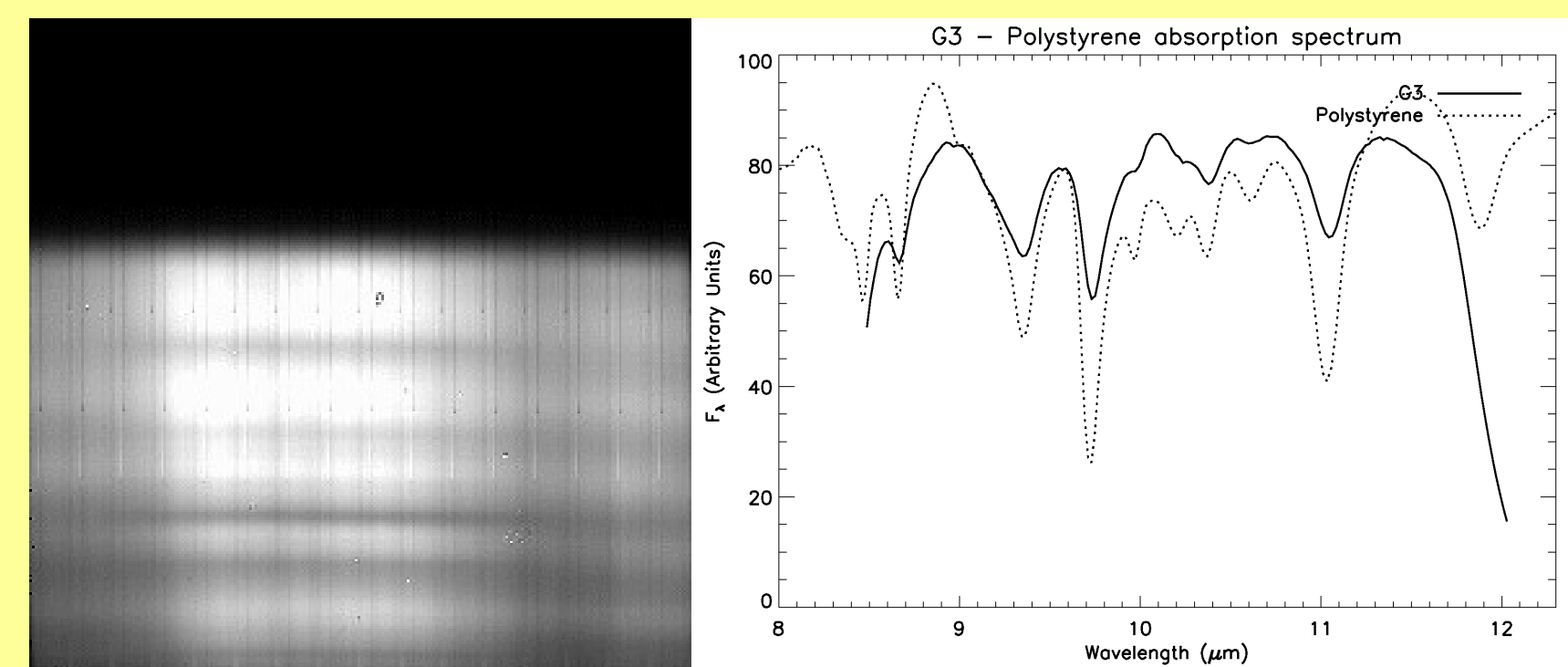
**Anti-Reflection Coating:** The entrance faces of Grisms 1 and 2 have been coated with a broad-band anti-reflection coating to reduce Fresnel losses and increase transmission. Grisms 3 and 4 were purchased with anti-reflection coatings already applied to the entrance faces.



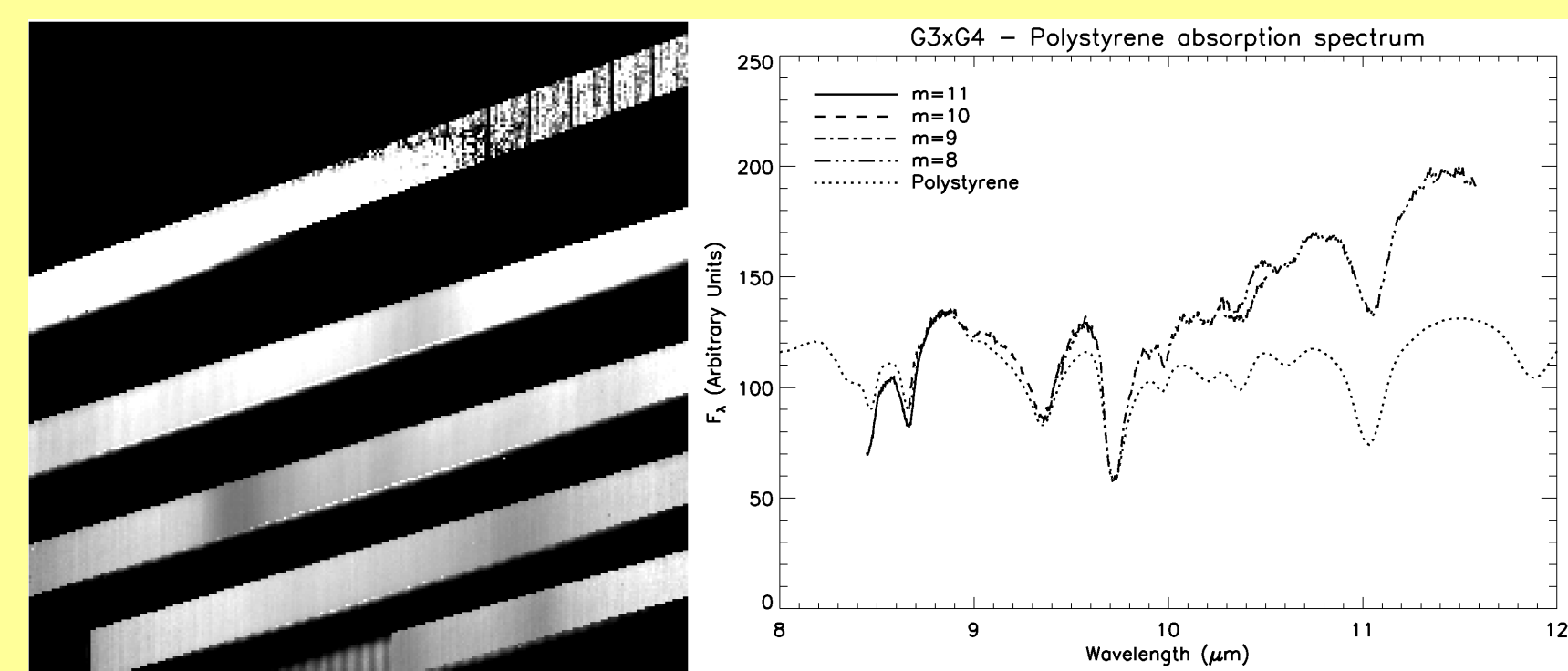
**Figure 5a:** Detector image and G1 extracted long slit spectrum of water vapor in emission above a liquid nitrogen background



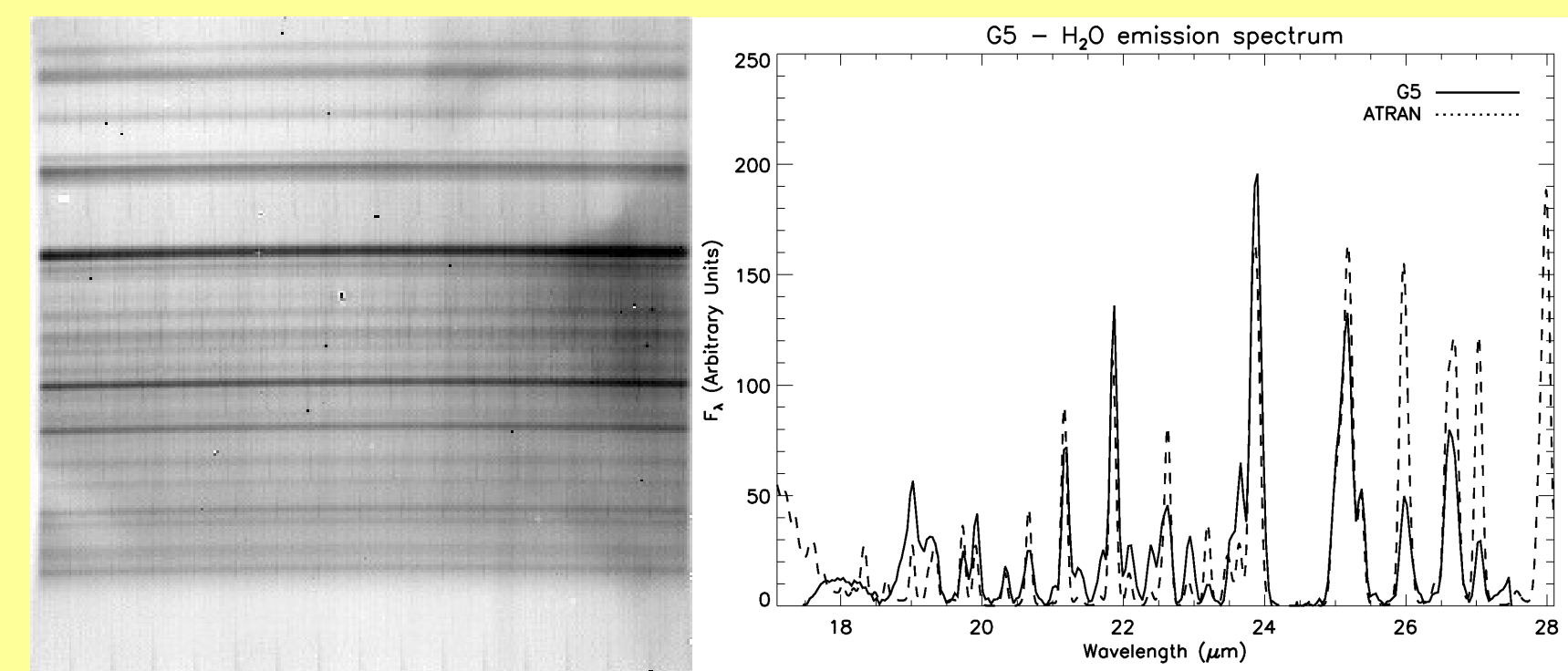
**Figure 5b:** Detector image and G1xG2 extracted cross dispersed spectrum of water vapor in emission over a liquid nitrogen background. Only the 18<sup>th</sup> order (highlighted in red) is extracted



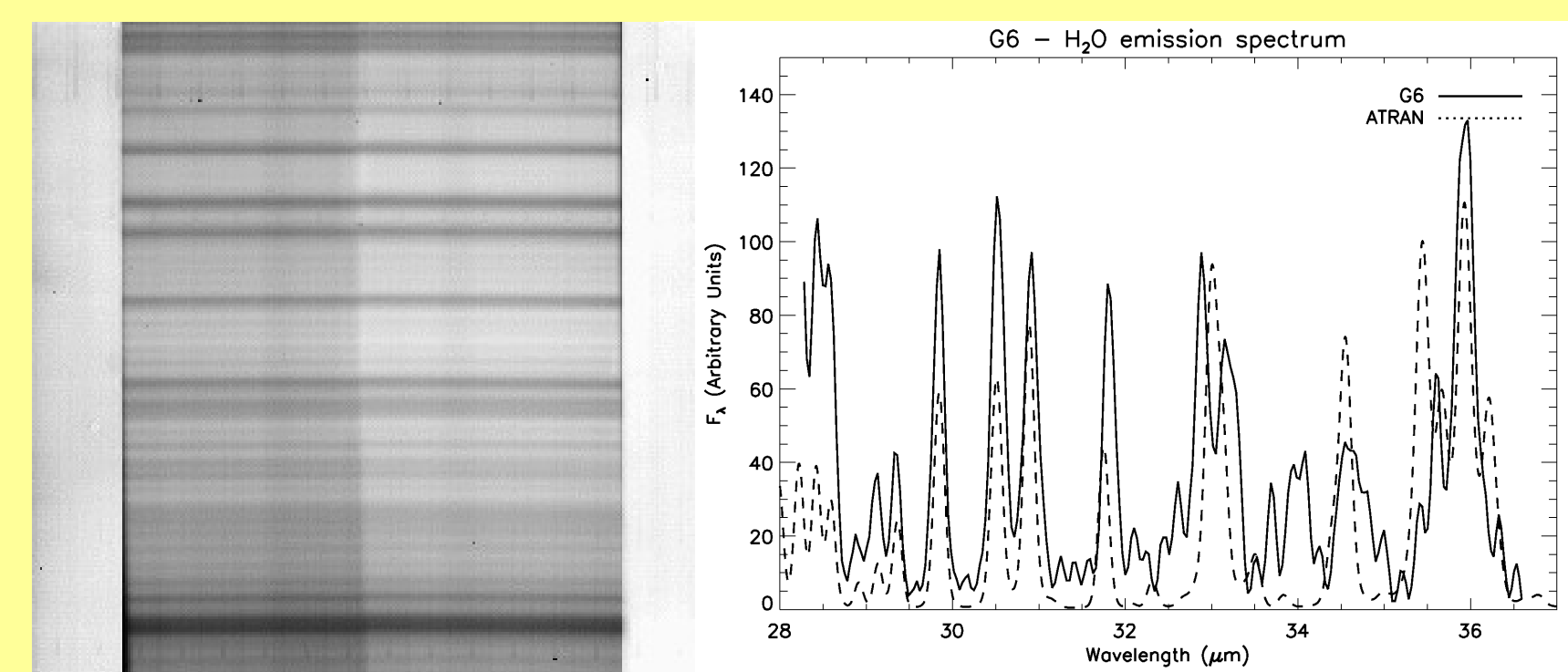
**Figure 5c:** Detector image and G3 extracted long slit spectrum of polystyrene film in absorption above a hotplate background



**Figure 5d:** Detector image and G3xG4 extracted cross-dispersed spectrum of polystyrene film in absorption above a hotplate background



**Figure 5e:** Detector image and G5 extracted spectrum of water vapor in emission above a liquid nitrogen background



**Figure 5f:** Detector image and G6 extracted spectrum of water vapor in emission above a liquid nitrogen background. G6 operates in second order, and was not used with the correct blocking filter in this exposure. The features at 33 and 36 microns are due to blue leaks in the filter showing up in third order.